

Two-dimensional electron density and temperature profiles of EUV light sources with 4.0% CE

Kentaro Tomita¹, Yuta Sato¹, Syoichi Tsukiyama¹, Raimu Fukada¹,
Fumitaka Ito¹, Kiichiro Uchino¹, Kouichiro Kouge², Tatsuya Yanagida²,
Hiroaki Tomuro², Yasunori Wada², Masahiro Kunishima²,
Takeshi Kodama², Hakaru Mizoguchi²



¹Kyushu University, Japan

²Gigaphoton Inc., Japan



Acknowledgements

Grate thanks for useful comments and discussions

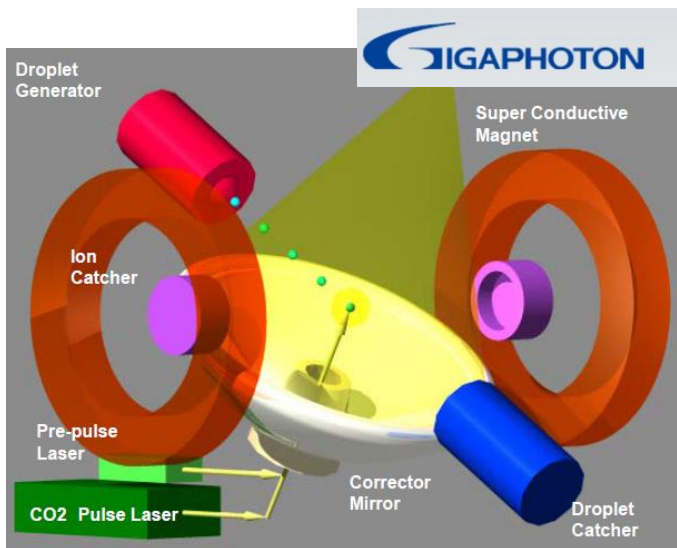
Prof. Katsunobu Nishihara
(Osaka Univ., Japan)



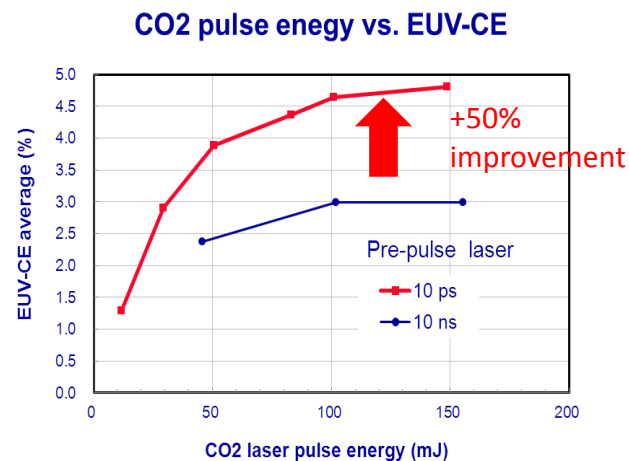
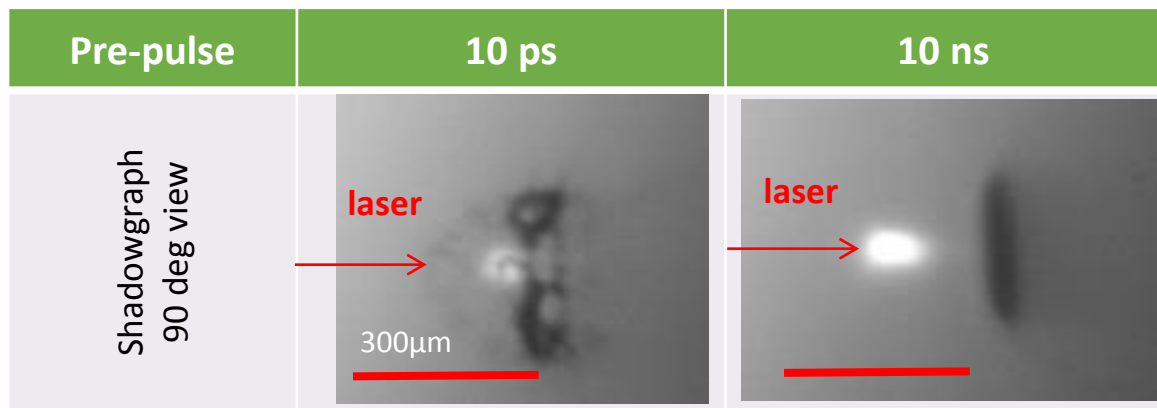
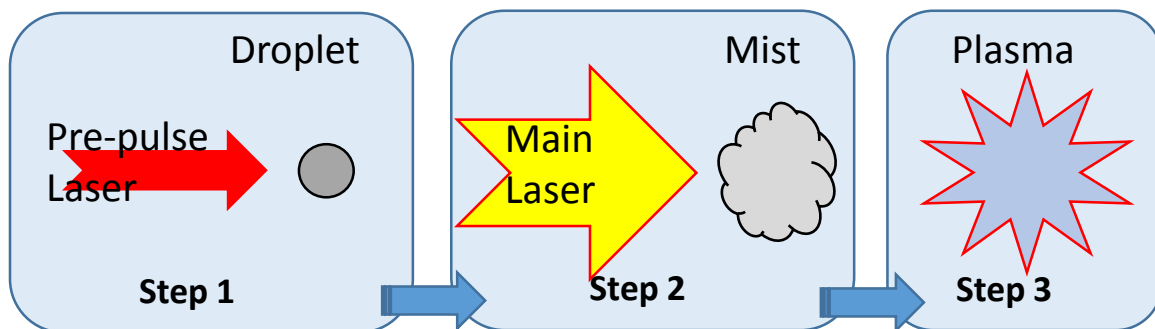
Dr. Atsushi Sunahara
(Purdue Univ., USA)



Motivation



Droplet Sn + Double pulse laser



↗ conversion efficiency (CE)

↗ output power

Structure?

Limitation ?

Improvement ?

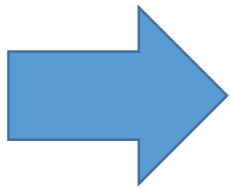
Diagnostic method

We have revealed the structure of EUV light source plasmas by using the world's best spatial resolution measurements.

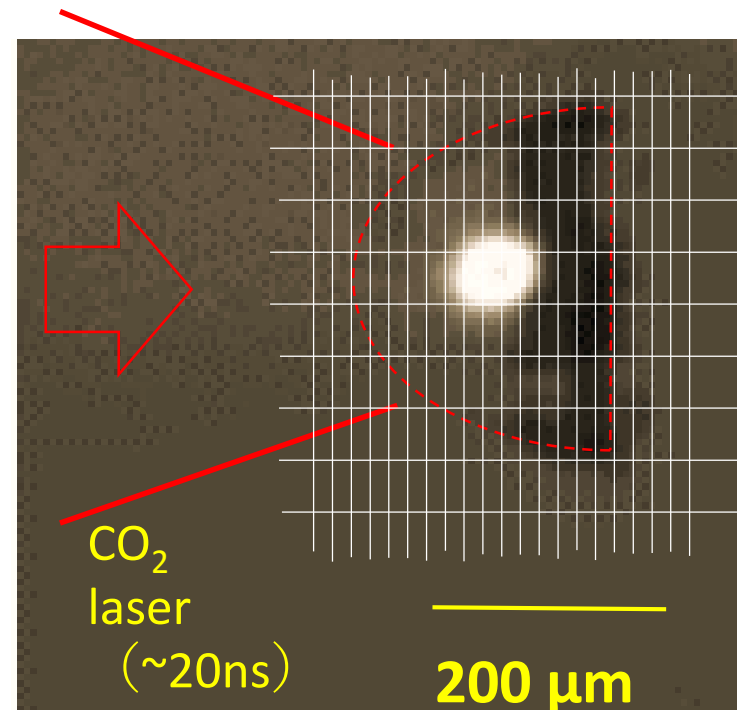
Collective Thomson scattering

- Spatial resolution ($\sim 20\ \mu\text{m}$)
- More than 200 points
- Electron density(n_e), temperature(T_e)

further information



Yuta Sato's poster



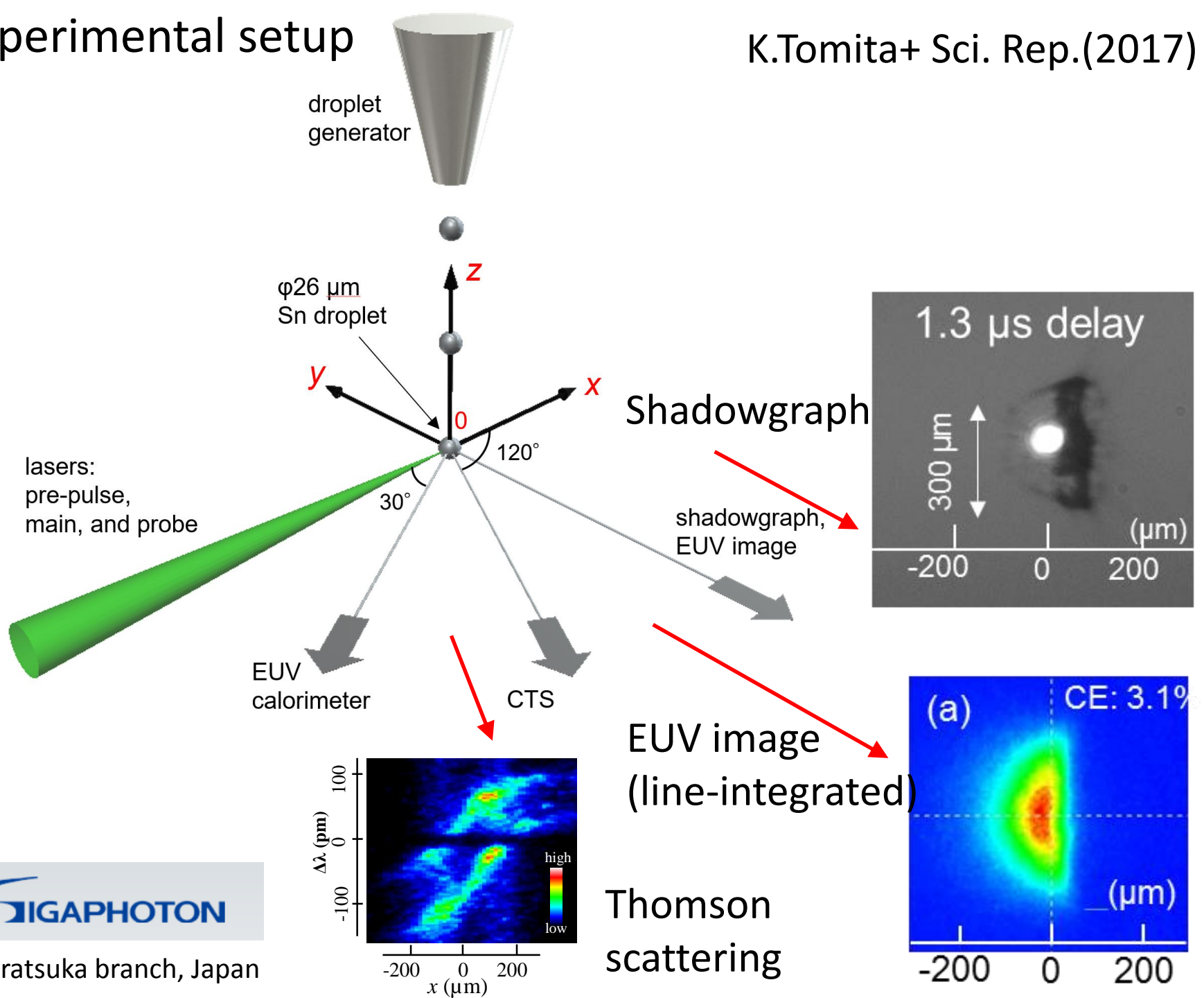
Shadow of expanded Sn target

Highlights

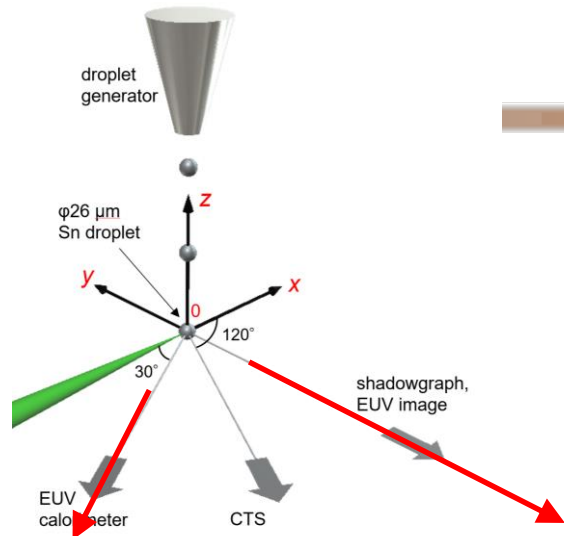
- We can clearly explain why the CEs were changed using the 2D-plasma parameters.
- It is possible to explain the CE values using the plasma parameters and atomic model.
- Monitoring of the 2D-plasma parameters can be a powerful tool for further improvement.

Experimental setup

K.Tomita+ Sci. Rep.(2017)



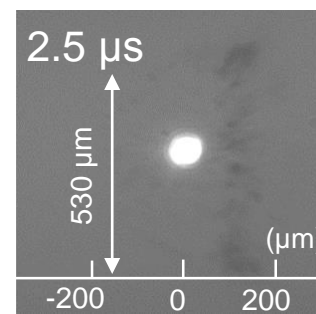
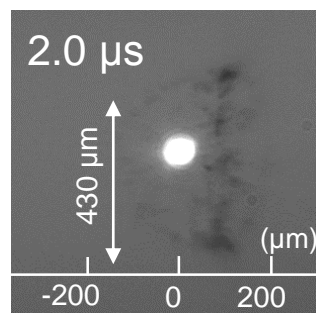
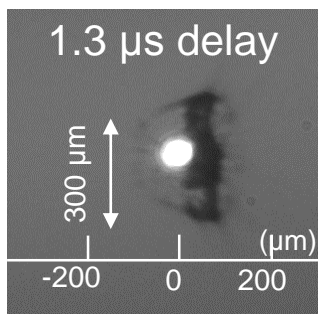
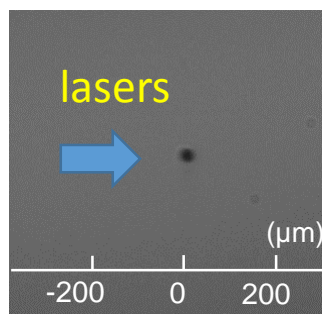
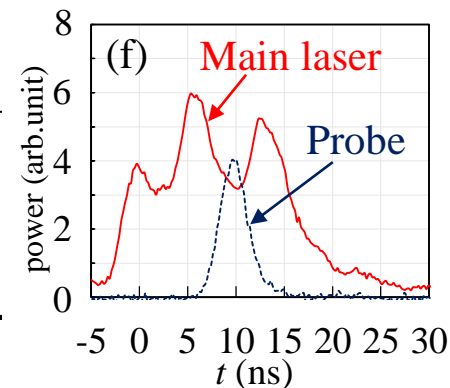
Plasma production scheme



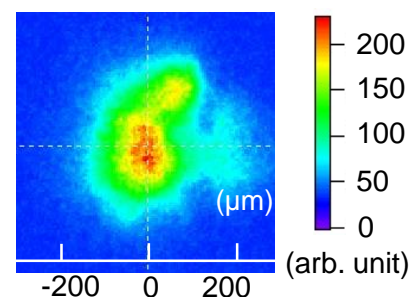
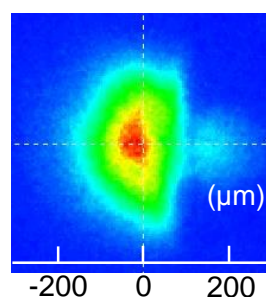
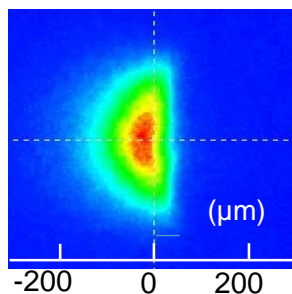
shadowgraph

Pre-pulse laser
(10 ps)

1.3-2.5 μs



EUV image

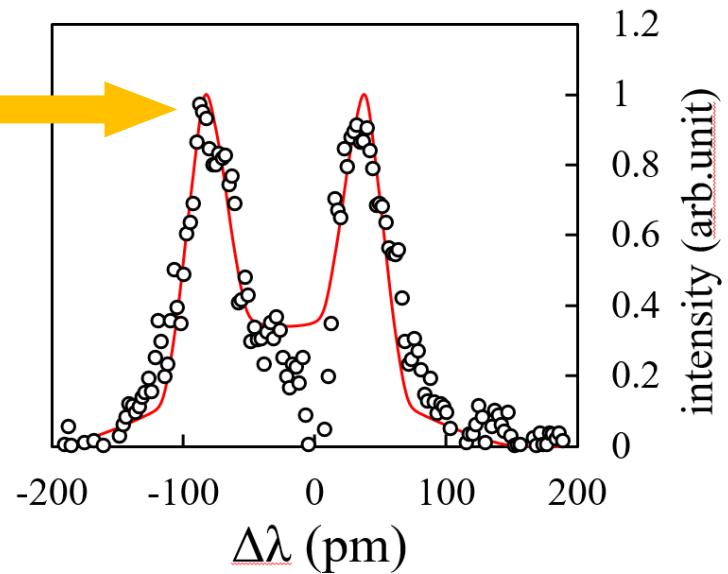
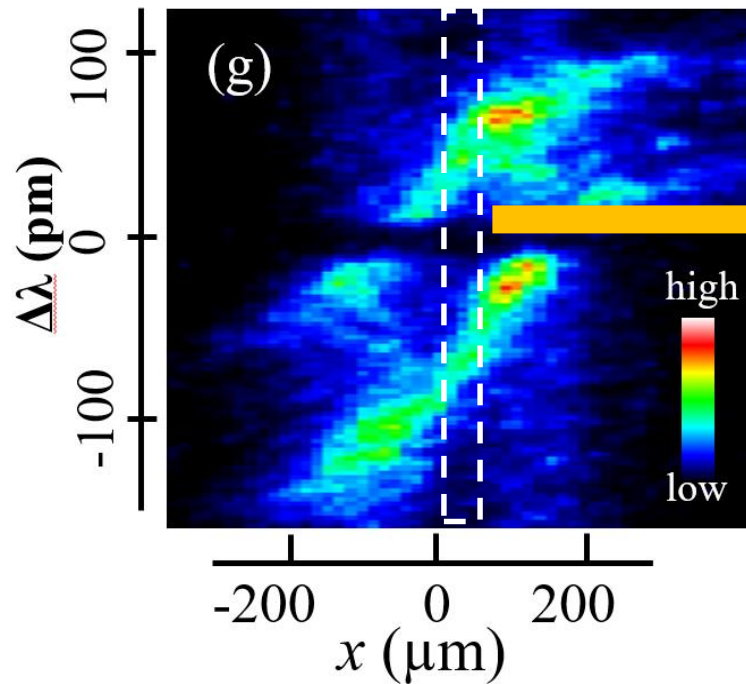
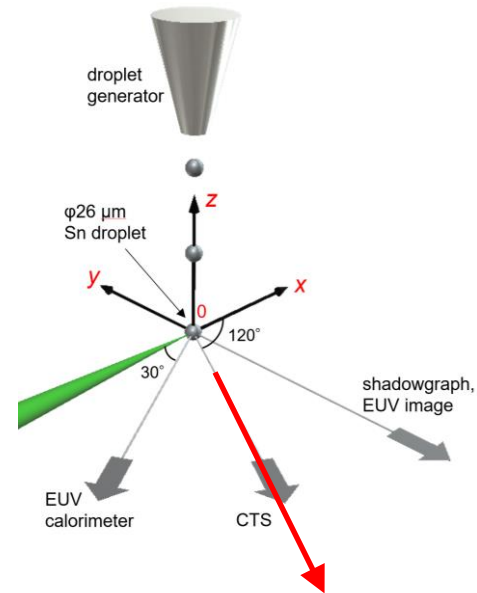
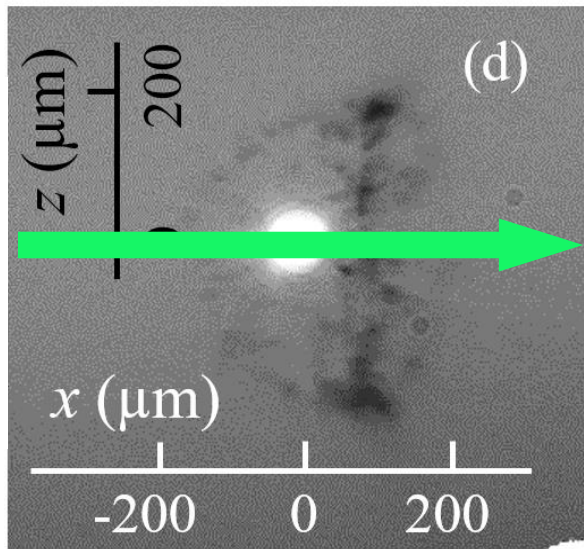


CE

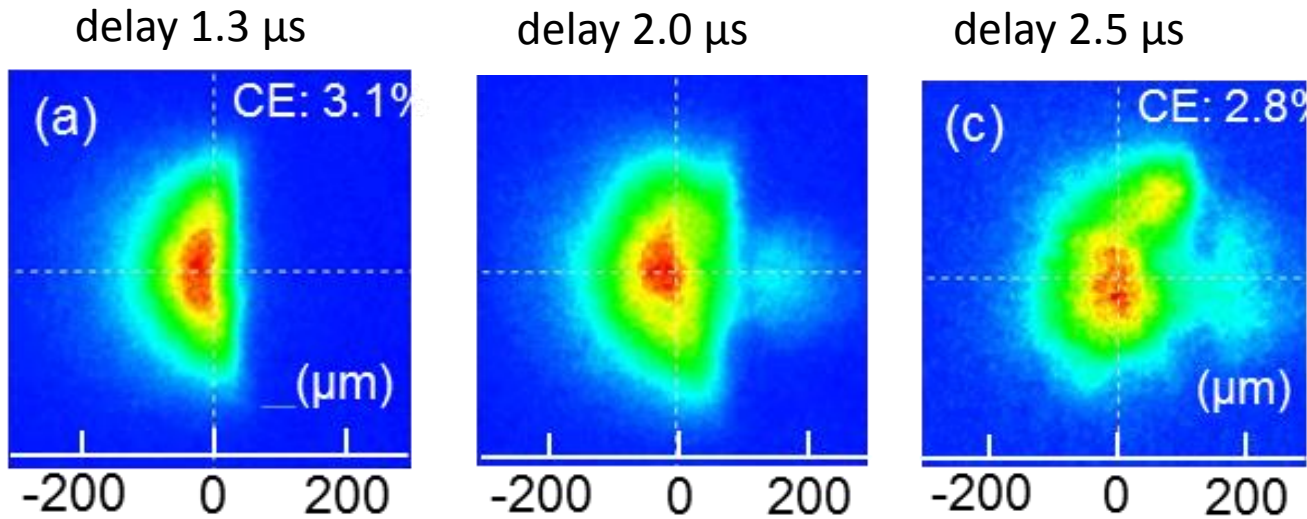
3.1 %

4.0 %

2.8 %

delay 2.0 μs Probe
laser

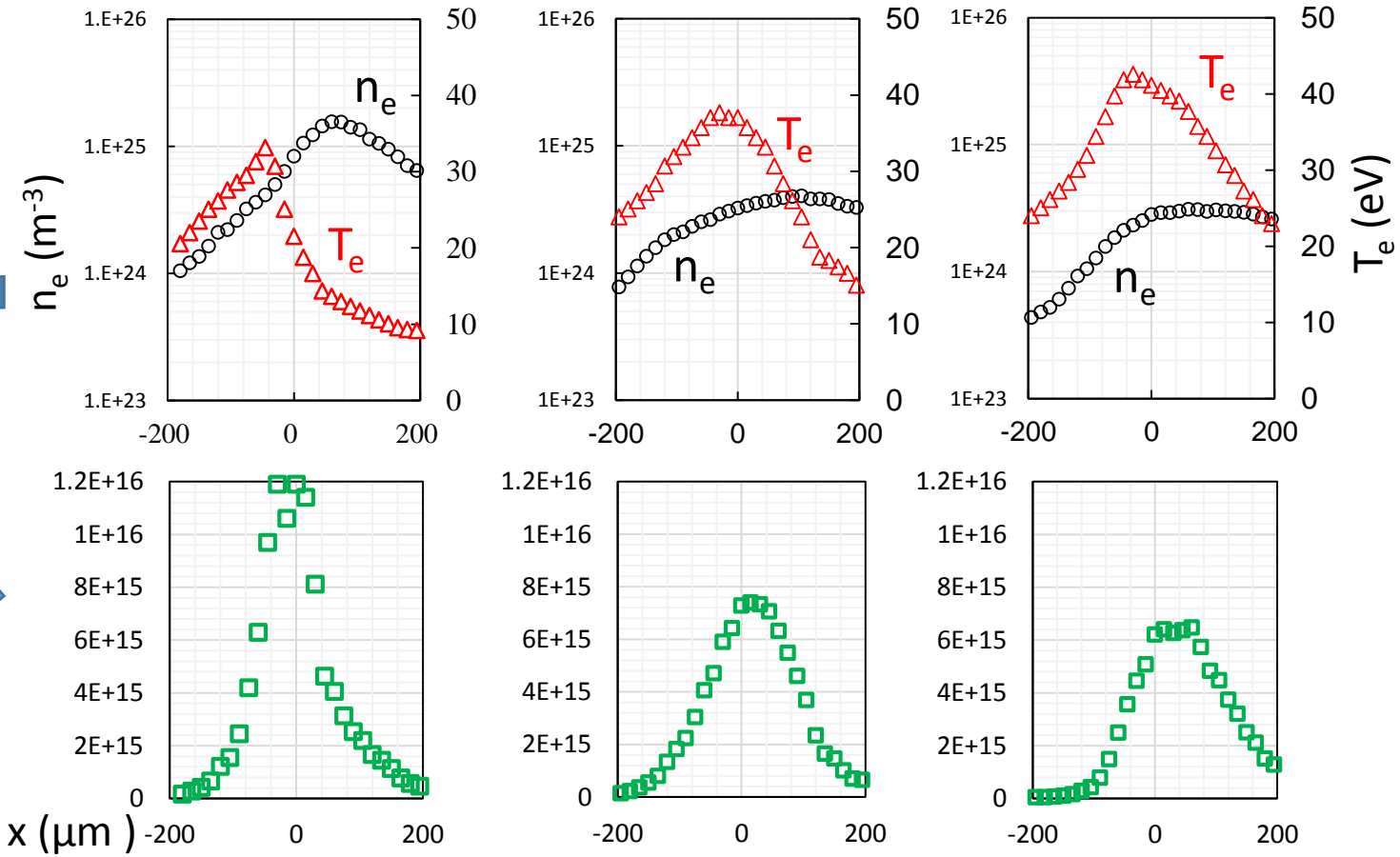
EUV image (line-integrated)



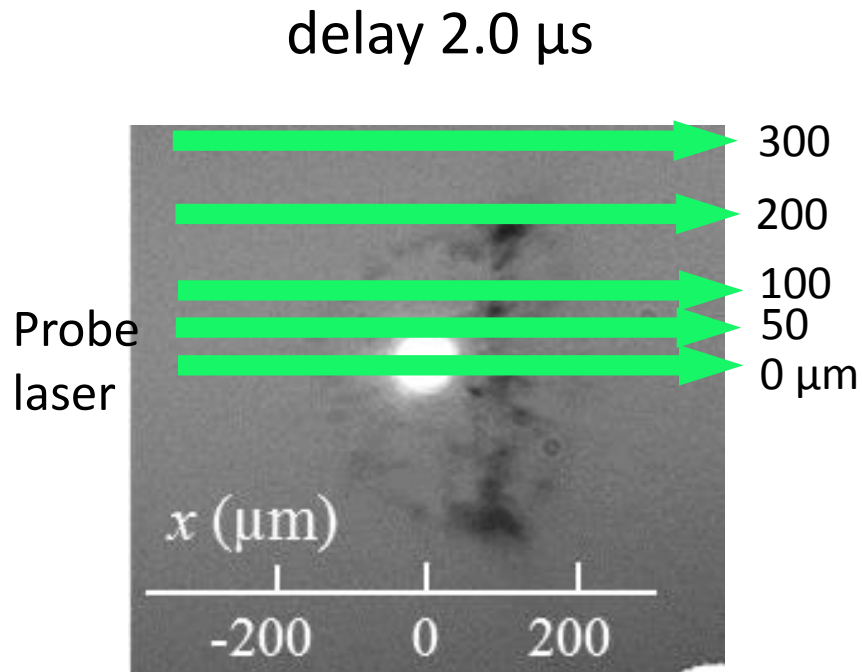
n_e and T_e

Hullac
code

Emissivity
($\text{W}/\text{m}^3/\text{eV}/\text{sr}$)



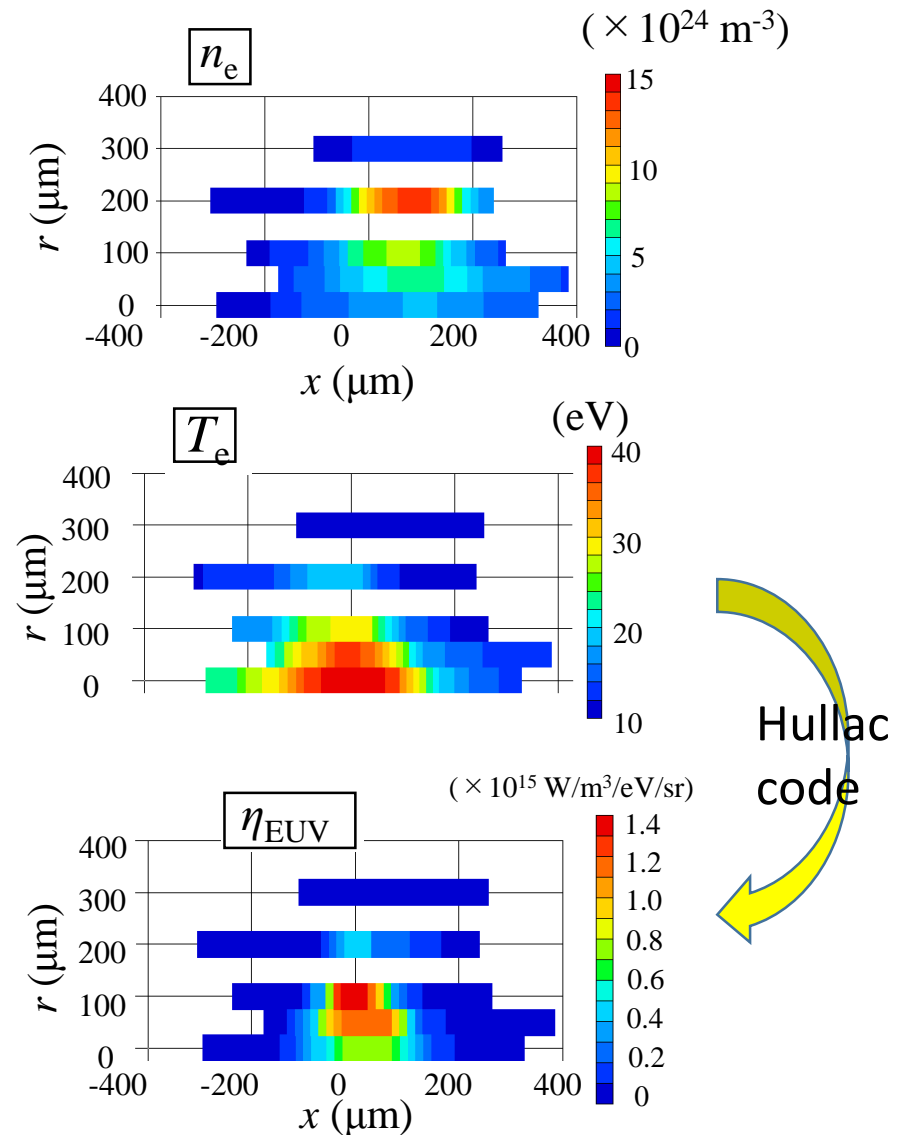
2D profiles of EUV source

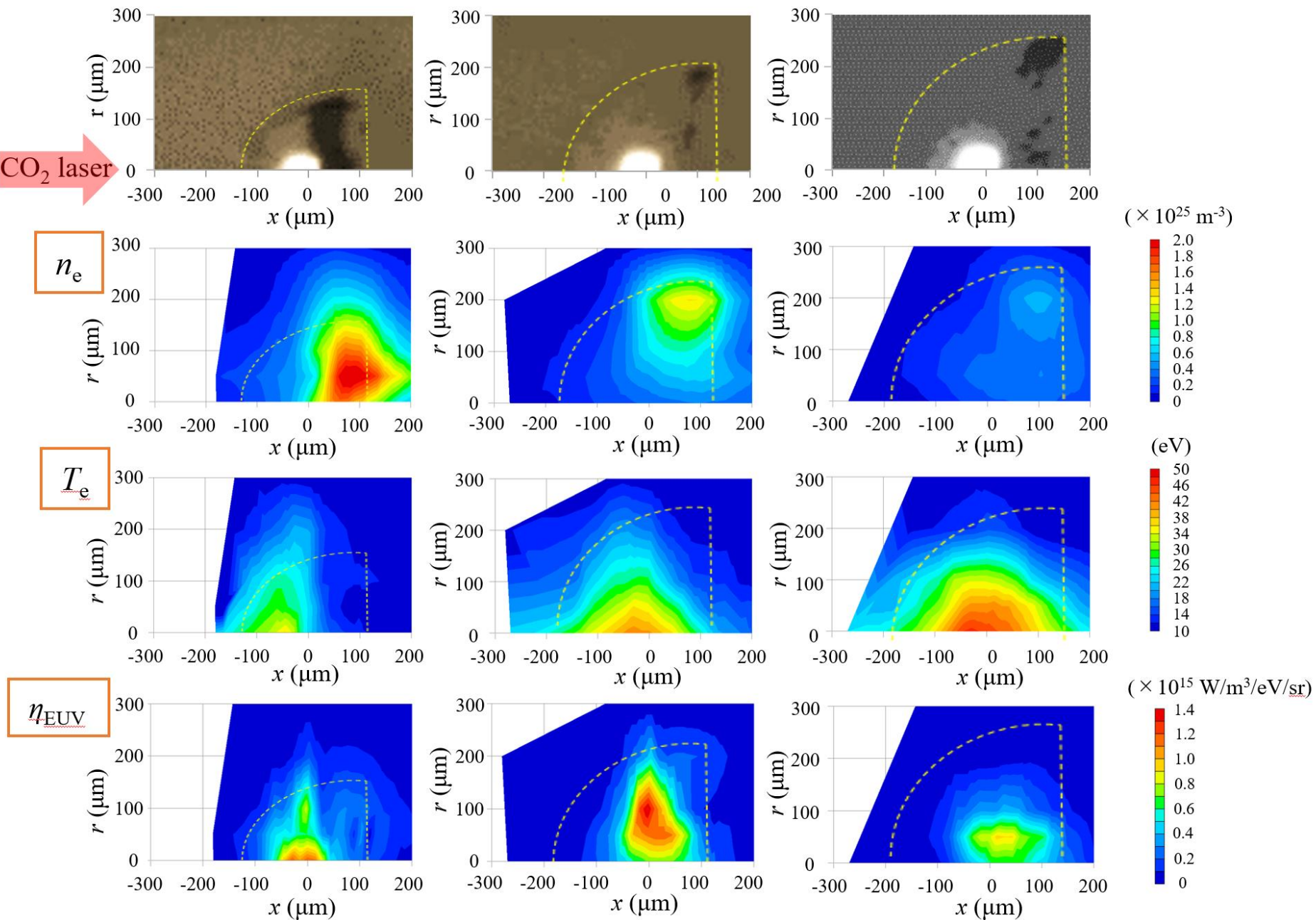


n_e : Hollow-like profile

T_e peak : center axis($r=0$)

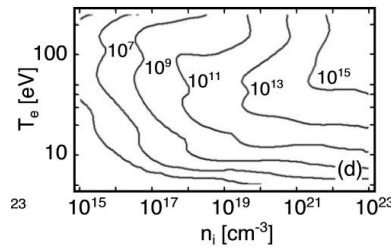
Emissivity peak: **not on $r=0$**



1.3 μs delay2.0 μs delay2.5 μs delayCO₂ laser

Discussion1

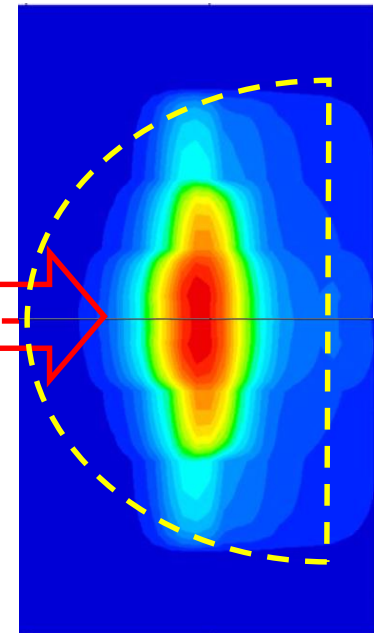
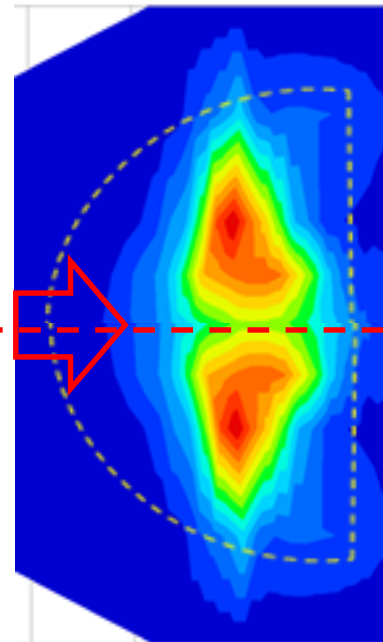
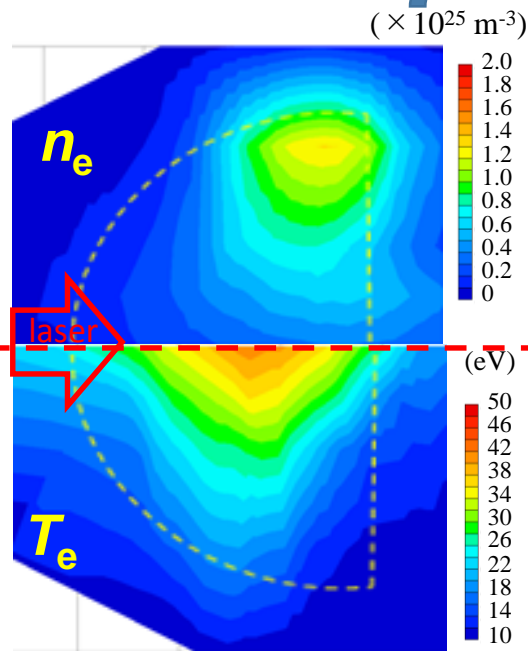
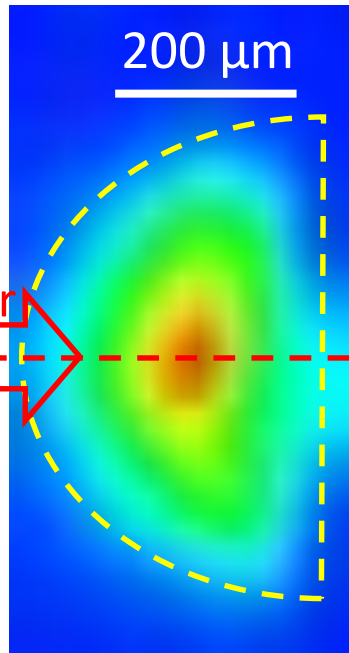
Atomic modeling

 n_e, T_e

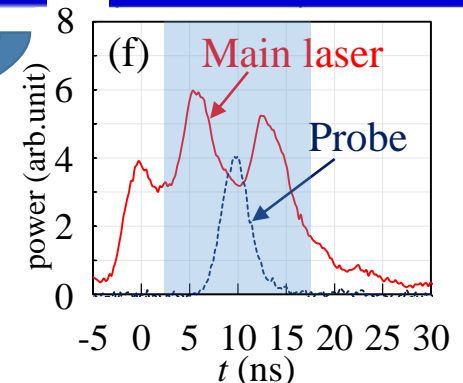
Emissivity

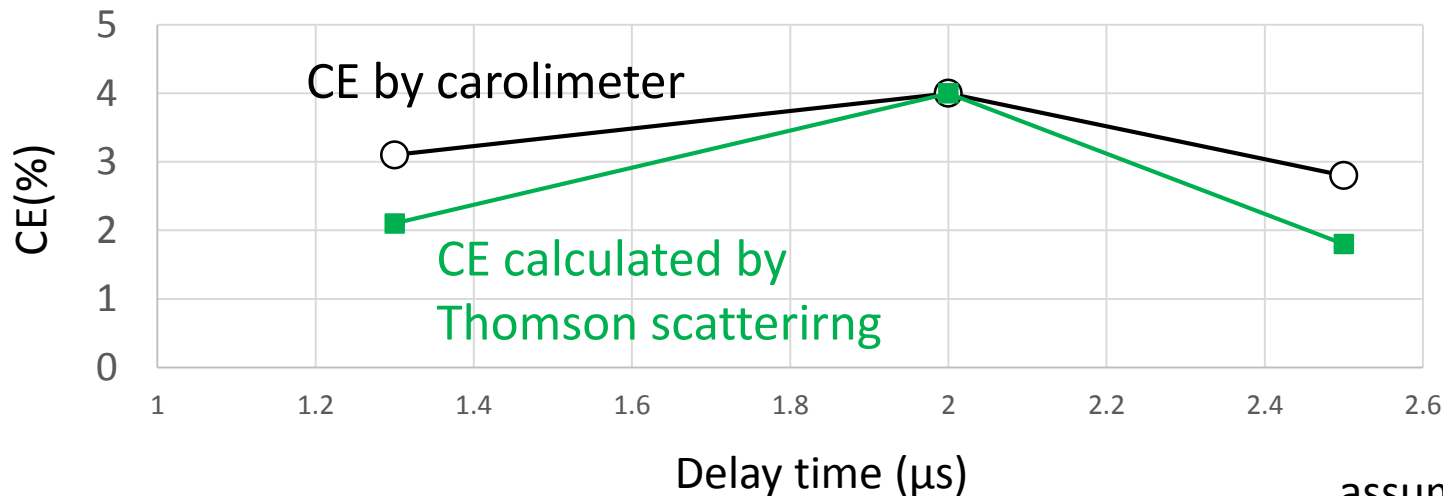
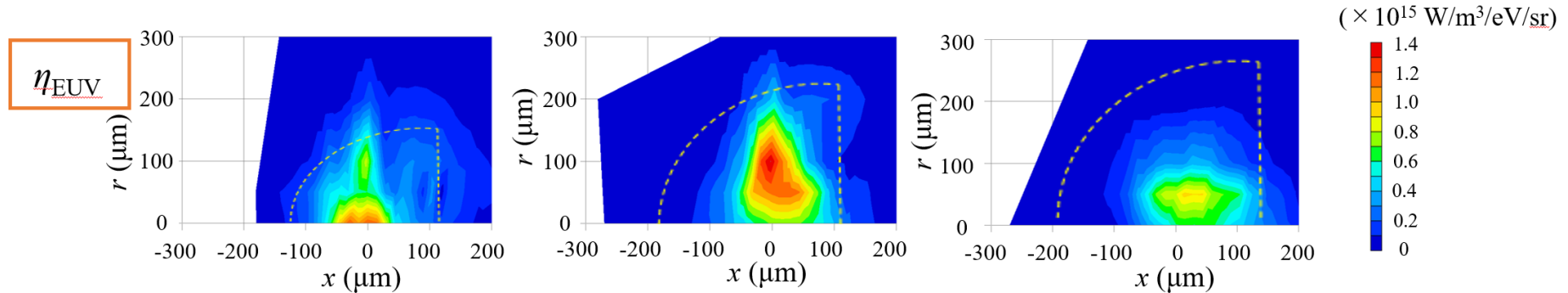
reconstructed image

EUV image



Results of line-integrated, self-absorbed



1.3 μs delay2.0 μs delay2.5 μs delay

assumptions:
axial symmetry
isotropic distribution
EUV duration 20ns

Conclusions

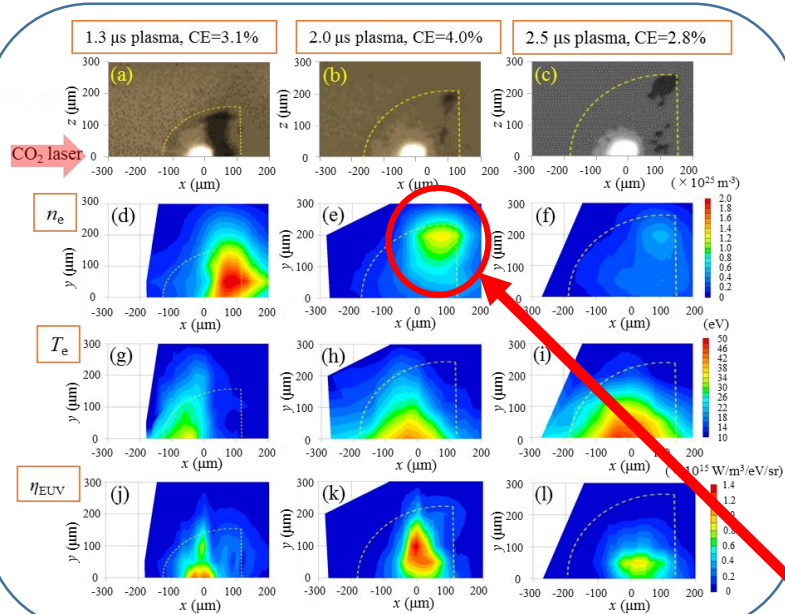


For the first time, we have revealed the detailed structure of EUV light source plasmas by the world's best spatial resolution Thomson scattering measurements.

Results clearly show that intense EUV were obtained where the plasma conditions are adequate, which are predicted by atomic modelling studies.

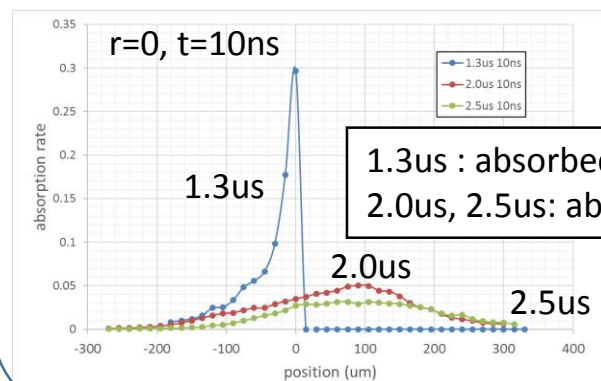
The 2D profile of the inband EUV emissivity (η_{EUV}) was theoretically calculated using the CTS results and atomic model (Hullac ode), which reproduced a directly measured EUV image reasonably well.

Future work: How can we improve the light source?



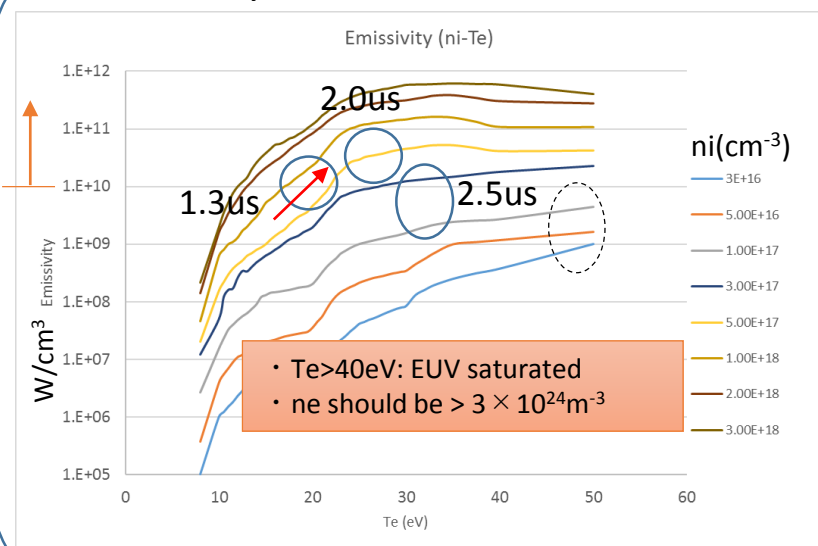
laser absorption (inverse bremsstrahlung)
is proportional to $n_e, T_e^{-3/2}$

$$k_{\text{abs}}^{\text{las}} = \frac{16\pi Z n_e^2 e^6 \ln \Lambda(\nu)}{3c \nu^2 (2\pi m_e k_B T_e)^{3/2} (1 - \nu_p^2 / \nu^2)^{1/2}} \quad (\text{Dawson})$$



1.3us : absorbed in narrow width
2.0us, 2.5us: absorbed in wide width

emissivity, ne, and Te (Sasaki+JAP2010)

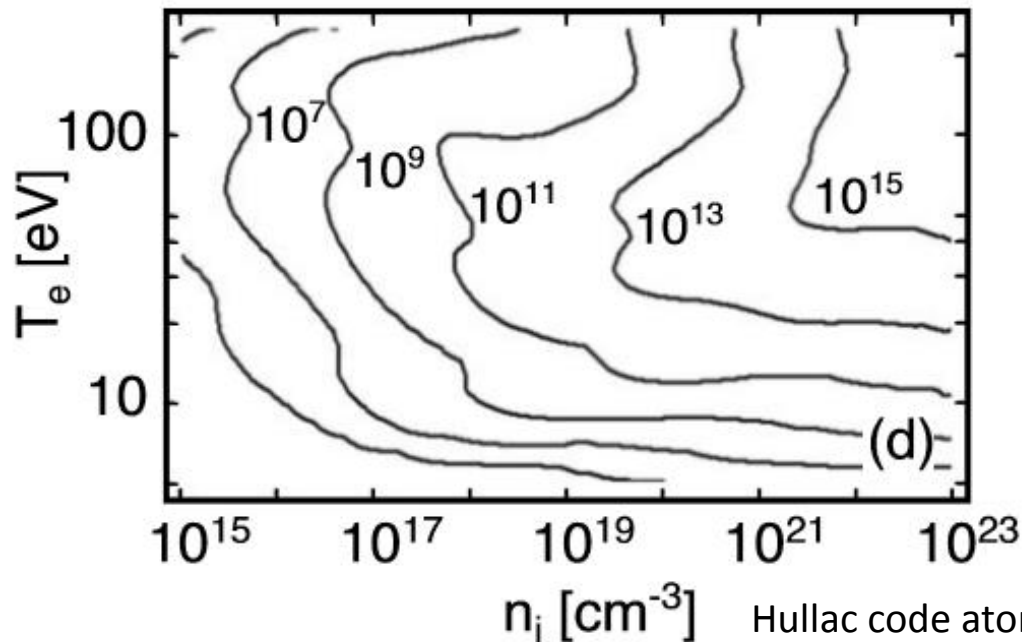


Even for the highest CE condition,
there are many low
temperature (<15eV) Sn ions within
the etendue limit!

Atomic modeling

Strong correlation between (T_e, n_e) and EUV emissivity

EUV emissivity(W/cm^3) contour diagram



Hullac code atomic modeling
A. Sasaki+ JAP(2010)

Optimum:

$$n_e = 10^{24} - 10^{25} \text{ m}^{-3}$$

$$T_e \sim 30 - 50 \text{ eV}$$

$$Z \sim 10$$

*non-linear dependence

e.g., 10eV, 30eV

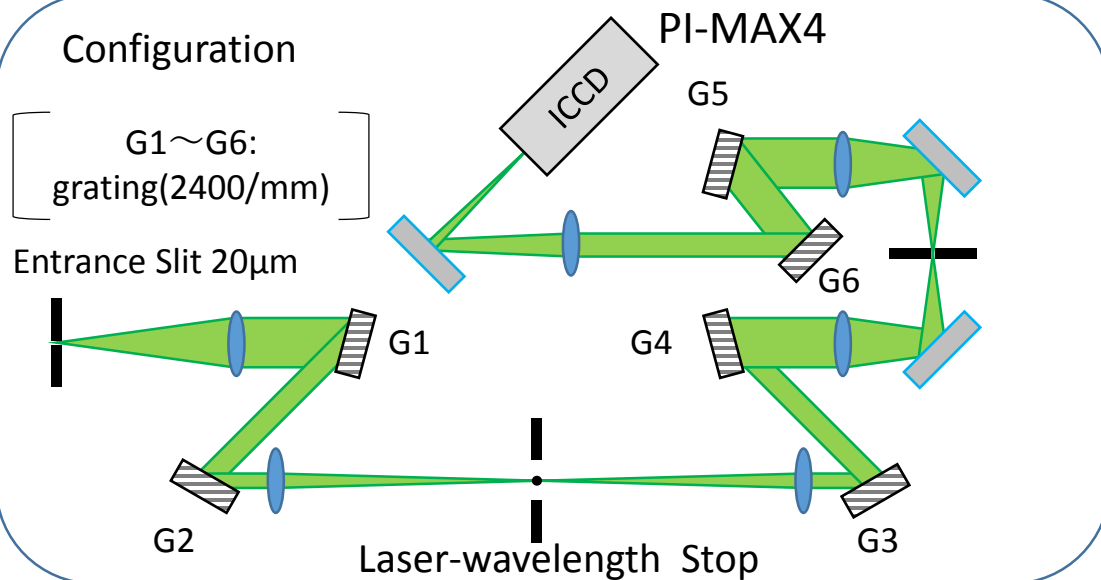
↓
 10^2 EUV

Plasma diagnostics are essential

K. Tomita+, APEX(2015)

Y. Sato+, JJAP(2017)

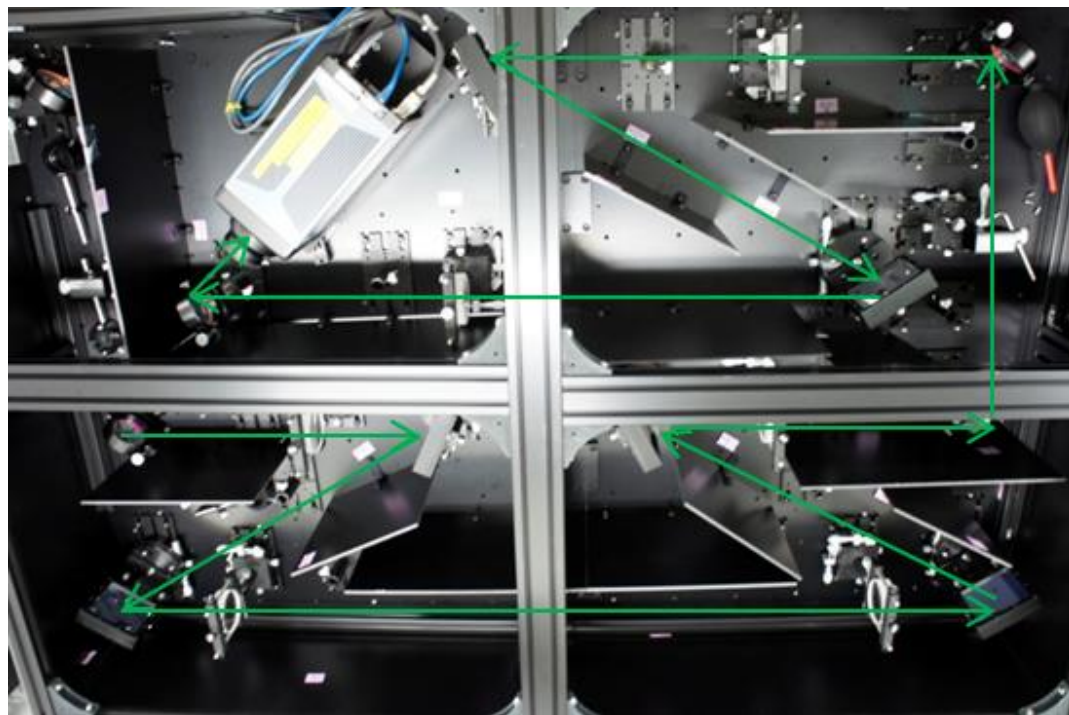
Configuration



spectrometer for CTS

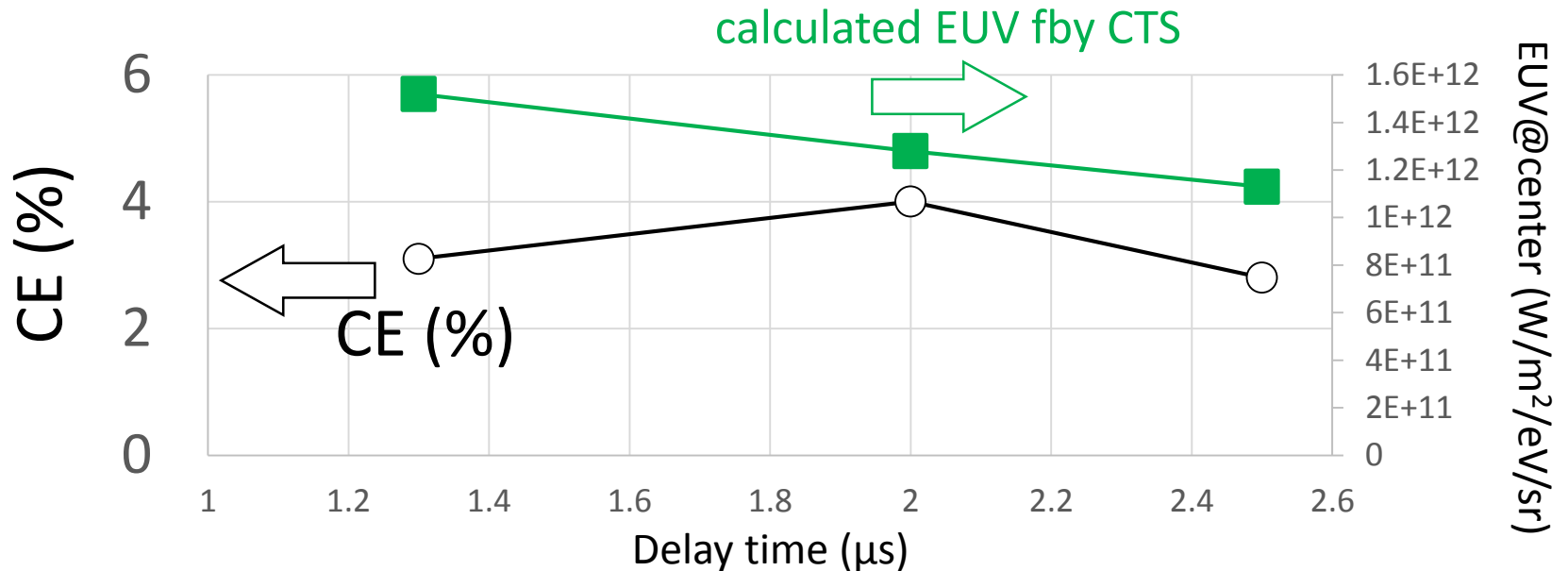
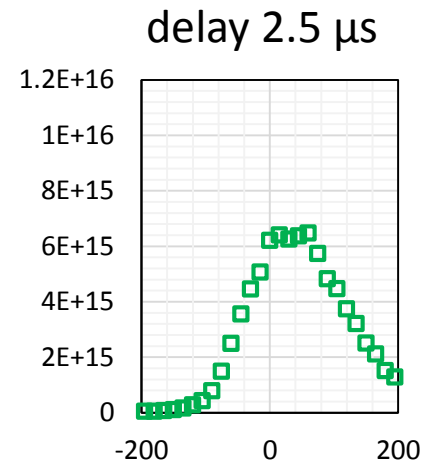
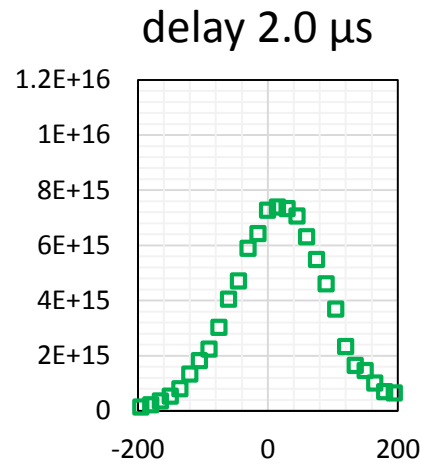
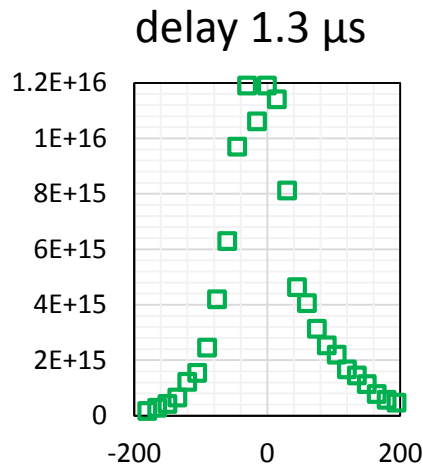
- spectral resolution (10pm)
- stray light rejection($\pm 14\text{pm}$)($\pm 0.5\text{cm}^{-1}$)

**a few tens of improvement
from conventional system**



2m

Emissivity
calculated by CTS
and atomic
modeling
(W/m³/eV/sr)



CE peak : delay **2.0 μ s**.

However, EUV peak @center axis : delay **1.3 μ s**